SECTION B - TECHNICAL PROPOSAL

Title of the project
Acronym
Acronym

The second direction of the project is dedicated to the research in the field of spatial dynamics, more precisely on turbulences and nonlinear phenomena which can appear in the motion of spatial objects (planets and satellites). The research will extend the previous results obtained by the team from University of Craiova in COMISIS and will bring new original contributions related to the development of new mathematical tools for investigating spatial nonlinear phenomena and to the design of new reduced-order multiple observers with applications to the flight of spacecraft. Important models arising from Gravitation can be analyzed following the specific procedure of nonlinear dynamics, the Ricci flow equation being a good example. Also, the research will focus on phenomena related to the motion of spatial objects and to the propagation of cosmic radiations; we shall concentrate our attention to the Lie group method. Because during the flight of spacecraft or satellites, their constrained and turbulent dynamics is characterized by different disturbances, the estimation of the states and/or unknown inputs is an important challenge; to reduce the number of sensors and measurements systems, we will design reduced-order multiple observers for Takagi-Sugeno nonlinear systems to estimate some of the dynamics' variables using intelligent algorithms and software. Beside the new architectures to be obtained, the software package is another expected impact of the project, in terms of specific delivrables.

On the third research direction, the relevance of Kelvin-Helmholtz instability (KHI) in the solar partially ionized plasma structures will be considered. The main expected impacts of the research will be related to a theoretical tool to predict the occurrence of unstable Kelvin-Helmholtz instabilities (dissipative instabilities or negative energy waves) in solar plasma with lower flow rates, as well as to the bulk transport properties of plasma through a statistical approach based on solving the stochastic differential equations. The existence of solar structures with partially ionized plasmas suggests the possibility to modify the criterion for the appearance of the KHI, allowing the onset of the instability. The existence of parameters whose values can be estimated only by statistically methods led us to propose an analytical and numerical method by which the classical two-fluid MHD formalism is modified by inserting stochastic analysis in determining the unstable solutions. The main goal is to obtain an equivalent instability condition containing parameters which may be compared to observations.

As we already mentioned, the research approach of COMASS will assume objectives which start from TRL 1 and end to TRL 4.